

## OPTICAL SATURATION

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### ABSTRACT

A general criterion is given for an amplifying medium to be self saturating. The concept of self-saturation gain has application to laser physics e.g. energy storage, and astrophysics (interstellar masers). Wide variations are shown to exist between different media.

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An amplifying medium reaches saturation when the gain coefficient is significantly decreased by stimulated emission. If the gain is sufficiently large this can take place via amplified spontaneous emission<sup>(1)</sup> and the medium then becomes self-saturating. This has important consequences for laser amplifiers<sup>(2)</sup> and astrophysical masers<sup>(3)</sup>.

The self-saturation gain  $G_s$  is most simply defined as the ratio of stimulated emission  $S$  at saturation to the spontaneous emission  $\mathcal{E}$  in the medium. Now  $S = B\Delta n I_s/c = \alpha I_s$ , where  $B$  is the Einstein coefficient for the stimulated process,  $\Delta n$  is the inverted population density,  $I_s$  is the saturation intensity,  $c$  the velocity of light and  $\alpha$  the small signal gain coefficient.

Therefore 
$$G_s^{-1} = \frac{\mathcal{E}}{\alpha I_s}$$

Because the inversion density and the gain coefficient are reduced by a factor 0.5 - 0.7 in the saturated case,  $G_s$  may be defined more precisely as the ratio of the saturation intensity  $I_s$  to the emissivity  $\mathcal{E}/\alpha$  of the medium. It should be noted that these expressions exclude solid angle because they represent an intrinsic property of the medium; geometric effects can be introduced later in solving specific problems via the equations of radiative transfer. The definition of  $G_s$  given above is perfectly general; as an example of its application a relation for  $G_s$  involving the rate constants of a two level system will now be derived.

For a radiative transition between two energy states with homogeneous broadening of the emission line we have<sup>(4)</sup>  $\epsilon = A_{21} n_2 h\nu_{21}$ , where  $A_{21}$  is the Einstein A coefficient connecting level 2 with level 1,  $n_2$  is the population of level 2,  $\nu_{21}$  is the radiation frequency;  $\alpha = A_{21} \lambda_{21}^2 \Delta n_{21} / \Delta\nu_{21}$ , where  $\lambda_{21} = \lambda_{21} / 2\pi$ ,  $\lambda_{21}$  is the wavelength of the radiation,  $\Delta\nu_{21}$  is the emission linewidth, the inversion density  $\Delta n_{21} = n_2 - \gamma_2 / \gamma_1 n_1$ ,  $\gamma_2$  and  $\gamma_1$  are the statistical weights of levels 2 and 1,  $n_1$  is the population of level 1; finally  $I_s = 2\Gamma_{21} h\nu_{21} \Delta\nu_{21} / A_{21} \lambda_{21}^2$ , where  $\Gamma_{21}^{-1} = \Gamma_2^{-1} + (\gamma_2 / \gamma_1) \Gamma_1^{-1} (1 - A_{21} \Gamma_2^{-1})$ , and  $\Gamma_1, \Gamma_2$  are the total decay rates - radiative plus non-radiative - of levels 1 and 2 connected to all other states, including the transition 2→1.

Using these relationships we find

$$G_s^{-1} = A_{21} n_2 / 2\Delta n_{21} \Gamma_{21}.$$

Solution of the rate equations<sup>(4)</sup> for levels 1 and 2 gives

$$\frac{n_2}{2\Delta n_{21}} = \frac{1}{2} \left[ 1 - \frac{\gamma_2}{\gamma_1} \left( \frac{A_{21}}{\Gamma_1} + \frac{W_1}{W_2} \frac{\Gamma_2}{\Gamma_1} \right) \right]^{-1}$$

where  $W_1, W_2$  are the excitation rates. Let  $q_{21} = n_2 / 2\Delta n_{21} = q(A, W, \Gamma)_{2,1}$ ,

therefore

$$G_s^{-1} = q_{21} A_{21} / \Gamma_{21} \quad (q_{21} \geq \frac{1}{2}).$$

A source of solid angle  $\Delta\Omega$  has a component of spontaneous emission  $\epsilon_1$  of magnitude  $\epsilon \Delta\Omega / 4\pi$ , if the spontaneous process is isotropic. An analogous gain saturation parameter can be defined in this case<sup>(3)</sup>

$$g_s^{-1} = \frac{\epsilon_1}{\alpha I_s} = q_{21} (A_{21}/\Gamma_{21}) (\Delta\Omega/4\pi) .$$

The following examples show that the magnitude of  $G_s$  can vary over a wide range.

For astrophysical masers values of  $G_s \sim 10^7$  to  $10^9$  were estimated previously<sup>(3)</sup>, whereas an ideal 4 - level laser has  $\Gamma_1 \gg \Gamma_2 = A_{21}$ , whence  $q = \frac{1}{2}$  and  $G_s = 2$ . Thus the gain in this case is essentially limited to  $\sim g_s \simeq 8\pi/\Delta\Omega$ . The minimum condition for an inversion  $\Delta n > 0$  is  $\gamma_1 \Gamma_1 > \gamma_2 A_{21}$ ; if  $W_1 = 0$ ,  $\Delta n_{21} \ll n_2$  and  $G_s \ll 1$  ( $\Gamma_2 \simeq A_{21}$ ). For nearly equal excitation rates  $W_2 - W_1 = \Delta W$  in a gas where  $\Gamma_2 = \Gamma_1 \gg A_{21}$  and  $\gamma_2 = \gamma_1$ ,  $G_s = (\Delta W/W_1)(\Gamma_2/A_{21})$ <sup>(5)</sup>. An infrared gas laser with  $W_2 \gg W_1$  and  $\Gamma_{21} \gg A_{21}$  has  $q_{21} \simeq \frac{1}{2}$  and  $G_s \simeq 2\Gamma_{21}/A_{21} \gg 1$ ; hence the high gain, high energy characteristic of this device<sup>(6)</sup>. For a 3 - level laser  $W_1 = -W_2$  and  $G_s = 2\Gamma_2/A_{21} \geq 2$ .

In the case of an inhomogeneously broadened emission line or a multi-line source, competition between stimulated emission on the different homogeneous components and cross-relaxation between them will increase the saturation gain  $G_s$  by a factor which is a function of the spectroscopic constants. This argument can be used to show that  $G_s > 55$  dB for the TEA CO<sub>2</sub> laser<sup>(7)</sup>; for Nd: glass  $G_s > 3$  dB.

Steady state conditions have been assumed throughout this analysis; in the transient case 'over-pumping' can produce temporary enhancement of pulse amplification<sup>(8)</sup>. In the absence of an external signal, the value of  $g_s$  gives a condition for the generation of saturated incoherent emission<sup>(1, 9)</sup>. The proposed 'super-radiant' X-ray laser<sup>(10)</sup> exceeds this requirement by several orders of magnitude.

There is also an important conclusion here for astrophysical masers; the self-saturation gain is always less than the threshold gain for maser oscillation via resonance scattering <sup>(11)</sup>. In these circumstances there would seem to be no meaningful distinction between the oscillator and partially saturated amplifier models of these <sup>(3, 11, 12, 13)</sup> sources

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