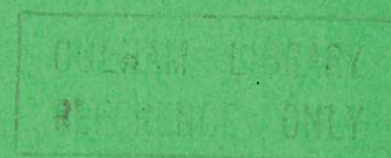
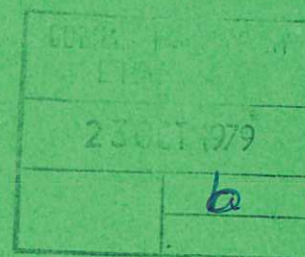




U K A E A

Report



THE EFFECT OF AN ANISOTROPIC CONDUCTING WALL ON THE INSTABILITIES OF A PINCH

D. C. ROBINSON

CULHAM LABORATORY
Abingdon Oxfordshire

1979

© - UNITED KINGDOM ATOMIC ENERGY AUTHORITY - 1979
Enquiries about copyright and reproduction should be addressed to the
Librarian, UKAEA, Culham Laboratory, Abingdon, Oxon. OX14 3DB,
England.

THE EFFECT OF AN ANISOTROPIC CONDUCTING WALL ON THE INSTABILITIES OF A PINCH

D C Robinson

Abstract

The dispersion relation for a thin skin pinch surrounded by an anisotropic thin conducting wall or liner and a perfectly conducting wall is obtained. Instabilities can grow on a timescale for the penetration of helical fields through the liner but near the marginal case for the liner they grow on a much faster hybrid timescale.

Introduction

The diffuse pinch which relies on a conducting wall to give it gross stability is always unstable if the conductivity of the wall is finite^[1]. We examine here for the simple case of a thin skin pinch the effects of a thin liner, whose conductivity is anisotropic, on the resultant growth rates.

Equilibrium

In a cylindrical coordinate system the equilibrium configuration is given by^[2]

$$\begin{aligned}\underline{B} &= B_{\theta 0}(0,0,b_i) \quad 0 < r < a \\ &= B_{\theta 0}(0,\frac{a}{r},b_e) \quad a < r < r_2\end{aligned}\quad (1)$$

where $B_{\theta 0}$ is the azimuthal magnetic field at the plasma surface ($r=a$). The pressure balance relation can be written as

$$\beta_{\theta} = 1 + b_e^2 - b_i^2 \quad (2)$$

where $\beta_{\theta} = 8\pi P/B_{\theta 0}^2$, b_i is the internal longitudinal magnetic field and b_e the external, which is negative for a reverse field pinch.

Equations

Perturbations of the form $\underline{\xi}(r)e^{\omega t + i(m\theta + kz)}$ are considered where $\underline{\xi}$ is the fluid displacement, m the azimuthal mode number and k the longitudinal wave number. For an incompressible plasma the perturbations of the magnetic field inside the plasma are given by

$$\underline{b} = ik b_i B_{\theta 0} \underline{\xi} \quad (3)$$

with the perturbed pressure and displacements given by

$$\begin{aligned}
 \tilde{p} &= \tilde{p}_0 I_m(kr) \\
 \xi_r &= - \frac{k \tilde{p}_0}{\rho_0 \omega^2} I_m'(kr) \\
 \xi_\theta &= - \frac{im}{r \rho_0 \omega^2} \tilde{p}_0 I_m(kr) \\
 \xi_z &= - \frac{ik}{\rho_0 \omega^2} \tilde{p}_0 I_m(kr)
 \end{aligned} \tag{4}$$

where I_m is the modified Bessel function and ρ_0 the plasma density. ($'$ denotes a derivative with respect to the argument of the function).

In the vacuum region between the plasma and the conducting wall at r_2 is a thin resistive liner of radius r_1 and thickness δ (Fig. 1). In the vacuum region the field perturbations, \underline{b} , are expressed in terms of the vector potential \underline{A} .

$$\begin{aligned}
 \underline{b} &= \nabla \wedge \underline{A}, \quad \underline{E} = - \omega \underline{A} \\
 A_z &= - A_z^1 K_m(kr) + A_z^2 I_m(kr) \\
 A_\theta &= \frac{m}{kr} A_z - i(A_\theta^1 I_m'(kr) - A_\theta^2 K_m'(kr)) \\
 b_z' &= k A_\theta - \frac{m}{r} A_z = - ik(A_\theta^1 I_m'(kr) - A_\theta^2 K_m'(kr))
 \end{aligned} \tag{5}$$

Dispersion relation

The continuity of total pressure across the plasma surface can be expressed in the form

$$- b_i B_{\theta 0} b_{zi} + b_e B_{\theta 0} b_{ze} + B_{\theta 0} b_{\theta e} - B_{\theta 0}^2 \frac{\xi_r}{a} = 4\pi \tilde{p}$$

which on using (3) and (4) becomes

$$- \left(\frac{k^2 b_i^2 B_{\theta 0}^2}{4\pi \rho_0 \omega^2} + 1 \right) \tilde{p}_0 I_m(kr) + \frac{B_{\theta 0}^2 k \tilde{p}_0}{4\pi a \rho_0 \omega^2} I_m'(kr) = - b_e \frac{B_{\theta 0}}{4\pi} b_{ze} - \frac{B_{\theta 0} b_{\theta e}}{4\pi} \tag{6}$$

At the plasma surface we have the condition on the electric fields

$$\underline{n} \wedge [\underline{E}] = - \omega \xi_r [\underline{B}] \tag{7}$$

where the bracket denotes the difference between the plasma and vacuum values at the interface and \underline{n} is the unit normal to the surface directed into the plasma. The z, θ components of (7) yield

$$\xi_r(a)(b_e + \frac{m}{ka}) B_{\theta 0} = i(A_{\theta}^1 I_m'(ka) - A_{\theta}^2 K_m'(ka)) \quad (8)$$

If we use the relation $b_{\theta e} = \frac{m}{ka} b_{ze}$ in (6) and combine (8) and (6) to eliminate \tilde{p}_0 we obtain

$$\frac{A_{\theta}^1 I_m(ka) - A_{\theta}^2 K_m(ka)}{A_{\theta}^1 I_m'(ka) - A_{\theta}^2 K_m'(ka)} = \left[\left(1 + \frac{k^2 b_i^2 B_{\theta 0}^2}{4\pi\rho_0 \omega^2} \right) \frac{I_m(ka)}{I_m'(ka)} \frac{4\pi\rho_0 \omega^2}{k^2 \left(\frac{m}{ka} + b_e \right)^2 B_{\theta 0}^2} - \frac{1}{ka \left(\frac{m}{ka} + b_e \right)^2} \right] \quad (9)$$

We now turn to the boundary conditions at the liner and conducting wall to determine $A_{\theta}^1, A_{\theta}^2$ to give a dispersion relation.

At the conducting wall the normal component of \underline{b} and the tangential components of \underline{E} are zero. If in the region $r_1 < r < r_2$, $A_z^3, A_z^4, A_{\theta}^3, A_{\theta}^4$ represent the coefficients of the vector potential as in (5) then

$$A_z^3 I_m(kr_2) = A_z^4 K_m(kr_2) \quad (10)$$

$$A_{\theta}^3 I_m'(kr_2) = A_{\theta}^4 K_m'(kr_2)$$

At the thin liner of thickness δ with conductivities $\sigma_z, \sigma_{\theta}$ the surface currents $j_{\theta} = -\omega\sigma_{\theta}A_{\theta}$, $j_z = -\omega\sigma_zA_z$ determine the jump in the magnetic field components. The continuity of the tangential components of the electric field expressed in the form

$$\underline{n} \wedge [\underline{E}] = 0$$

allows us to relate $A_{\theta}^3, A_{\theta}^4$ to $A_{\theta}^1, A_{\theta}^2$

$$A_{\theta}^3 = \frac{[A_{\theta}^1 I_m'(kr_1) - A_{\theta}^2 K_m'(kr_1)]}{I_m'(kr_1)K_m'(kr_2) - K_m'(kr_1)I_m'(kr_2)} \cdot K_m'(kr_2) \quad (11)$$

with a similar expression for A_{θ}^4 .

Using the expressions for the magnetic field components b_z , b_θ across the liner, (5) and (11) we obtain

$$\frac{A_\theta^1 I_m(kr_1) - A_\theta^2 K_m(kr_1)}{A_\theta^1 I_m'(kr_1) - A_\theta^2 K_m'(kr_1)} = -\frac{\omega \delta \sigma_z k r_1^2}{m^2 + \frac{\sigma_z}{\sigma_\theta} k^2 r_1^2} + \frac{C_2(r_1, r_2)}{C_1(r_1, r_2)} \quad (12)$$

where $C_1(r_1, r_2) = I_m'(kr_1) K_m'(kr_2) - K_m'(kr_1) I_m'(kr_2)$

and $C_2(r_1, r_2) = K_m'(kr_2) I_m(kr_1) - I_m'(kr_2) K_m(kr_1)$.

From (12) some general conclusions can be drawn. For $m \geq 1$ and long wavelength instabilities, $kr_1 \ll 1$ the growth rate will only depend on σ_z . For $m = 0$ the rate depends on σ_θ and for $m > 1$ but short wavelength, $k^2 r_1^2 \sigma_z / \sigma_\theta \gg 1$ it also depends on σ_θ . We can now obtain the final dispersion relation from (9) and (12) by eliminating A_θ^1 , A_θ^2 . After some algebra this is

$$\begin{aligned} & -\frac{1}{ka(\frac{m}{ka} + b_e)^2} + \left(\frac{1 + k^2 b_i^2 B_{\theta 0}^2}{4\pi \rho_0 \omega^2} \right) \frac{I_m(ka)}{I_m'(ka)} \cdot \frac{4\pi \rho_0 \omega^2}{k^2(\frac{m}{ka} + b_e)^2 B_{\theta 0}^2} \\ & = \frac{C_2(a, r_2) + \frac{\omega \delta \sigma_z k^2 r_1^3}{m^2 + \frac{\sigma_z}{\sigma_\theta} k^2 r_1^2} \cdot C_2(a, r_1) C_1(r_1, r_2)}{C_1(a, r_2) + \frac{\omega \delta \sigma_z k^2 r_1^3}{m^2 + \frac{\sigma_z}{\sigma_\theta} k^2 r_1^2} C_1(a, r_1) C_1(r_1, r_2)} \quad (13) \end{aligned}$$

Solution

If $\delta = 0$, i.e. no liner, this expression reduces to that given by Tayler^[2] (other dispersion relations similar to (13) have been given for different configurations^[3,4,5]). The plasma is stable if and only if the above cubic dispersion relation has no root with a positive real part. The stability criterion is of course not dependent on the presence of the liner. Stability results for values of b_e , b_i and r_2/a with no liner are given in the literature^[2].

Introducing the notation $\Gamma^2 = 4\pi \rho_0 \frac{a^2 \omega^2}{B_{\theta 0}^2}$, $S = \frac{\tau_w}{\tau_{A\theta}}$, $X = ka$, $\gamma = \sigma_z / \sigma_\theta$

with $\tau_w = 2\pi r_1 \delta \sigma_z$, the penetration time through the liner for a poloidal

field, and $\tau_{A0} = \frac{a\sqrt{4\pi\rho_0}}{B_{00}}$ the dispersion relation can be written

$$\Gamma^2 \frac{I_m(x)}{I_m'(x)} + x^2 b_i^2 \frac{I_m(x)}{I_m'(x)} = x + (m + x b_e)^2 \cdot \frac{C_2(a, r_2) + h C_2(a, r_1) C_1(r_1, r_2)}{C_1(a, r_2) + h C_1(a, r_1) C_1(r_1, r_2)} \quad (14)$$

$$\text{with } h = \frac{2\Gamma S k^2 r_1^2}{m^2 + \gamma k^2 r_1^2}$$

or in a more explicit form

$$\Gamma^3 + \frac{\Gamma^2}{\alpha} + \Gamma T(r_1) + \frac{T(r_2)}{\alpha} = 0 \quad (15)$$

$$\text{where } T(r_2) = x^2 b_i^2 - x \frac{I_m'(x)}{I_m(x)} - (m + x b_e)^2 \frac{I_m'(x)}{I_m(x)} \cdot \frac{C_2(a, r_2)}{C_1(a, r_1)}$$

- if equated to zero is the marginal stability condition at the wall. $T(r_1) = 0$ is the condition at the liner, and

$$\alpha = \frac{2S k^2 r_1^2}{m^2 + \gamma k^2 r_1^2} \cdot \frac{C_1(a, r_1) C_1(r_1, r_2)}{C_1(a, r_2)} \quad (16)$$

If $T(r_2) > 0$ the plasma is stable. The situation which is of most interest is $T(r_2) < 0$ but $T(r_1) > 0$ i.e. the plasma would be stable if the liner were perfectly conducting. If the value of S is large then to a first approximation

$$\Gamma \approx - \frac{T(r_2)}{T(r_1)} \cdot \frac{(m^2 + \gamma k^2 r_1^2) \cdot C_1(a, r_2)}{2S k^2 r_1^2 C_1(a, r_1) C_1(r_1, r_2)} > 0 \quad (17)$$

i.e. the growth time is proportional to τ_w but multiplied by a complicated function depending on the k value of the instability and the anisotropy of the liner, γ . A bellows liner with a convolution factor of C has $\gamma \sim C^{-2}$, consequently all instabilities with $m > 1$ and $kr_1 < C$ are unaffected by the convolution factor. For the particular thin skin model of the pinch used here all long wavelength instabilities are stable i.e. $T(r_1), T(r_2) > 0$ for kr_1, kr_2 small so that no limiting expression can be obtained from (17).

If the conducting wall is removed i.e. $r_2 \rightarrow \infty$ then (17) can be simplified somewhat as

$$\frac{C_1(a_1, r_2)}{C_1(r_1, r_2)} \rightarrow \frac{K_m'(ka)}{K_m'(kr_1)}$$

which for $m > 1$ and ka small gives $(r_1/a)^{m+1}$ and Γ reduces to the form

$$\Gamma \approx - \frac{T(r_2)}{T(r_1)} \cdot \frac{m}{S(1 - a^{2m}/r_1^{2m})} \quad (17A)$$

i.e. the growth time is essentially the field penetration time through a thin wall in the presence of an internal conductor of radius a but reduced by the m number. Equation (17) is a rather more general expression for the penetration of helical fields through a thin walled vessel for which there are expressions in the literature [6,7].

For $m = 0$ with $r_2 \rightarrow \infty$ and in the particular limit that $ka \rightarrow 0$ we obtain

$$\Gamma \approx - \frac{T(r_2)}{T(r_1)} \frac{\gamma}{2S} K_1(kr_1) I_1(kr_1) \quad (17B)$$

i.e. the growth time is $\sim 4\pi\sigma_\theta r_1 \delta K_1(kr_1) I_1(kr_1)$ which gives the usual penetration time $2\pi\sigma_\theta r_1 \delta$ if kr_1 is small but with σ_θ the effective component. If kr_1 is large then the time is much shorter $\sim 2\pi\sigma_\theta r_1 \delta / kr_1$.

A notable exception to the expected behaviour given in (17) is when $T(r_1) \rightarrow 0$ and an instability just fails to grow at a rate $\tau_{A\theta}^{-1}$. In this case

$$\Gamma \approx \frac{3\sqrt{T(r_2)}}{\alpha^{1/3}} \propto \frac{1}{S^{1/3}} \quad (18)$$

or $\omega \sim \tau_w^{-1/3} \tau_{A\theta}^{-2/3}$ as noted previously [8]

Computations

The dispersion relation (14) has been solved for a number of cases of interest. Fig. 2 shows the resultant growth rate as a function of ka . The region of instability in this and most other cases is exemplified by $ka \sim 1$, as is indeed the case for quite general diffuse pinch configurations. In this case $b_i = 0.9$, $b_e = -0.5$ so that $\beta_\theta = 0.44$, $r_1/a = 1.3$, $r_2/a = 1.7$

and $T(r_2) < 0$, $T(r_1) > 0$. The value of S is taken to be 100 which is typical of diffuse pinch experiments with a liner such as Zeta. Curve (a) gives the growth rate as a function of ka for the case of no liner. With a liner present having $\sigma_z/\sigma_\theta = 8, 4, 2, 1, \frac{1}{2}$ the growth rate is progressively reduced as shown by the curves (b) to (f) and as can be seen from equation (17). In most experiments with a liner $\sigma_z/\sigma_\theta < 1$ so that no further reduction in the growth rate of modes with $m > 1$ occurs as σ_θ is progressively increased (a reduction in growth rate would occur for $m=0$ modes). Consequently the copper rings present around the liner in the Zeta device^[9] would not be expected to convey better stability properties on the plasma unless modes with $m=0$ were important. Note that even though $S = 100$ the reduction in growth rate associated with the presence of the liner is only tenfold.

Fig. 3 demonstrates the reduction in growth rate with increasing S for the same parameters as Fig. 2., Γ is proportional to S^{-1} as expected from (17) but if b_e is such that $T(r_1) \rightarrow 0$, i.e. $b_e = - .553$, then the liner is not very effective in reducing the growth rate as indicated by (18). For $b_e = - .553$ we do indeed find that Γ is proportional to $S^{-1/3}$.

Even for the optimistic case of $\Gamma \propto S^{-1}$, a pinch operating with $n\tau_E \geq 10^{15}$ cm⁻³ sec at a density of 10^{14} cm⁻³ would require a 65 cm thick copper shell at a radius of 3m to avoid an $m=1$ instability for one energy confinement time. At a density of 10^{15} cm⁻³ and a radius of 1.2 m the Cu thickness falls to 17 cms. These results indicate that some slow feedback control of the $m=1$ ideal magnetohydrodynamic instability and $m=1$ tearing mode would be necessary.

Conclusions

The introduction of a corrugated liner into an otherwise unstable diffuse pinch can reduce the growth rate significantly to a value somewhat larger than that associated with the penetration of a helical field through a thin resistive wall. If the anisotropy of the wall conductivity - $\frac{\sigma_z}{\sigma_\theta} < 1$ then there is little effect on the growth rate of modes with $m \geq 1$ but the growth rate for $m = 0$ is reduced. If the plasma is close to marginal stability with respect to the liner then the instability grows on a hybrid timescale of the penetration time and Alfvén transit time.

Acknowledgement

I would like to thank Adrian Furzer for solving the dispersion relation and obtaining the results shown in Figs. 2 and 3.

References

- [1] D Pfirsch, H Tasso Nuclear Fusion, 1971, vol.11, p.259.
- [2] R J Tayler 1957, Proc. Phys. Soc. B70, p.1049.
- [3] M D Kruskal et al. Physics Fluids 1958, vol.1, p.421.
- [4] J P Goedbloed, D Pfirsch, H Tasso Nuclear Fusion, 1972, vol.12, p.649.
- [5] M Tanaka et al. Nuclear Fusion, 1973, vol.13, p.119.
- [6] V S Mukhovatov, V D Shafranov, Nuclear Fusion, 1971, vol.11, p.605.
- [7] V F Aleksin, S S Romanov, Zhurnal Tech. Fiz. 1974, vol.44, p.1877.
- [8] A S Furzer, D C Robinson, 7th European Conference on Controlled Fusion and Plasma Physics, Lausanne, 1975, vol.1, p.114.
- [9] J T D Mitchell et al. Proc. Inst. Elec. Eng. A106, Supplement 2 (1959) 101.

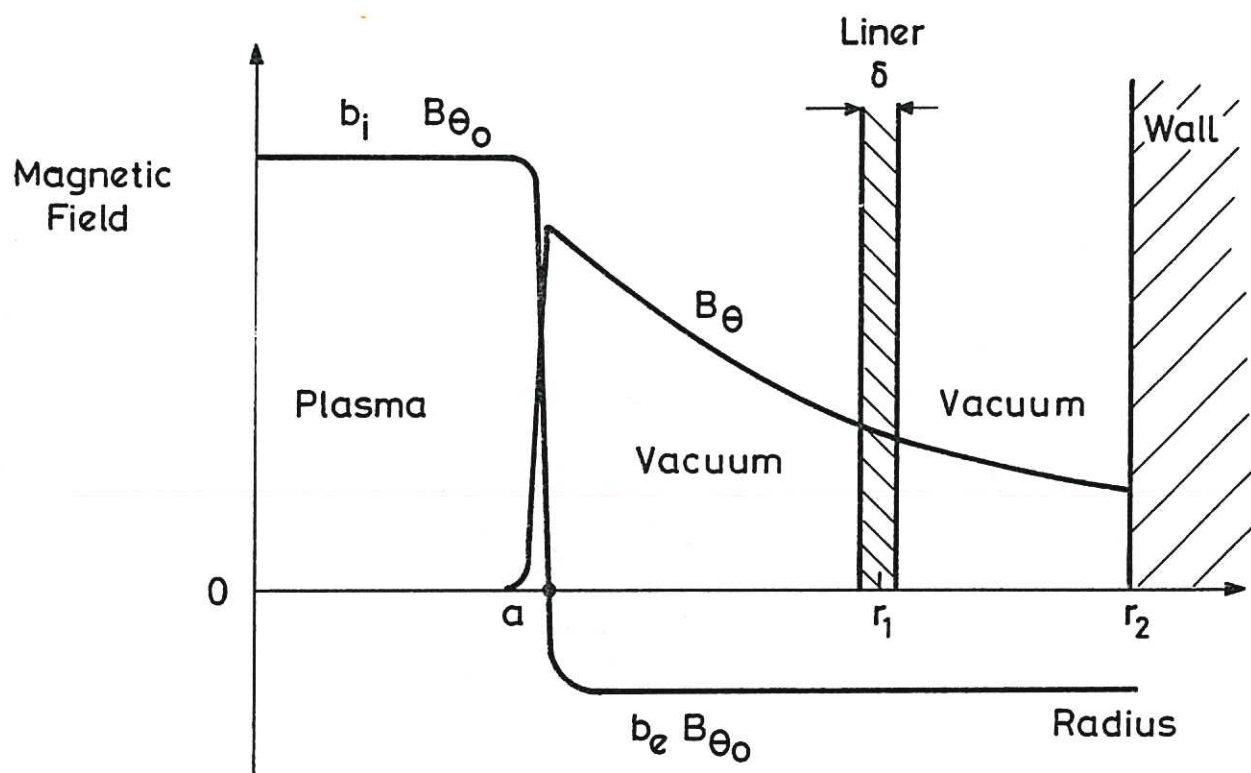


Fig. 1 Equilibrium field configuration showing edge of plasma, liner and conducting wall.

$$\frac{r_1}{a} = 1.3$$

$$\frac{r_2}{a} = 1.7$$

$$S = 100$$

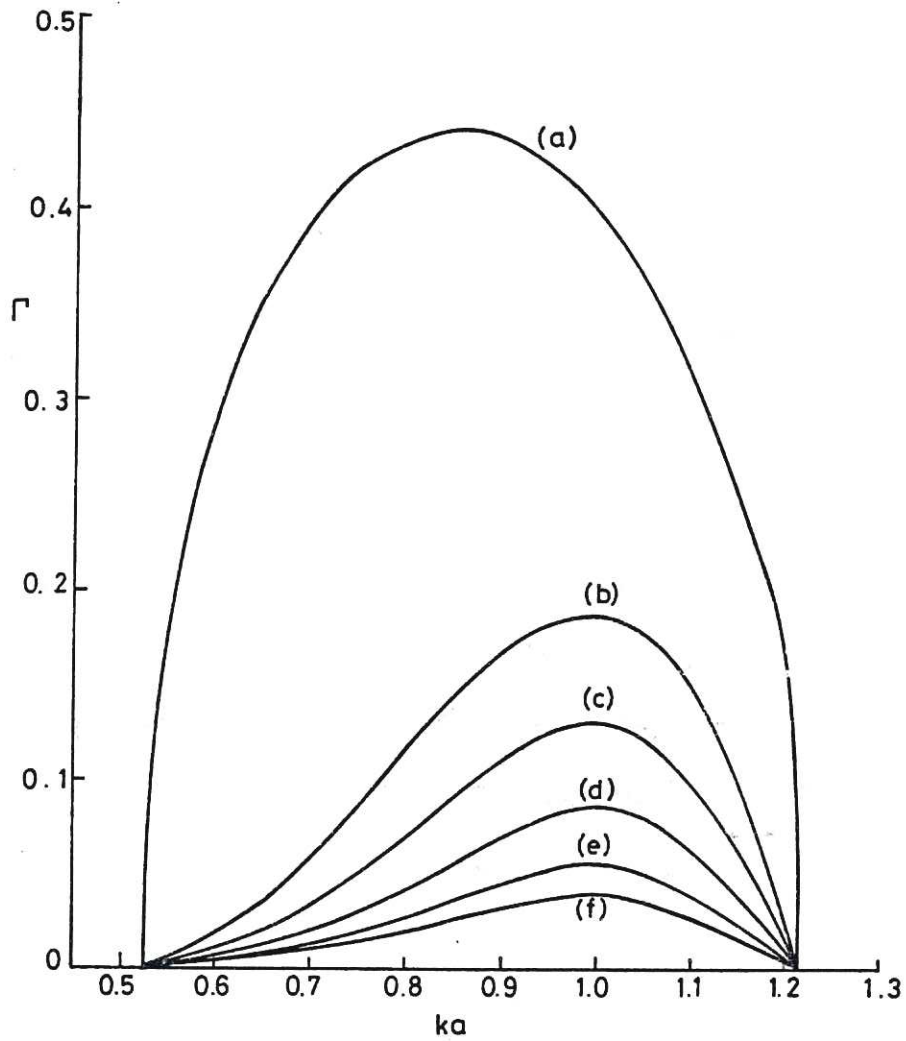


Fig. 2 Normalised growth rate as a function of ka for
 (a) no liner (b) $\frac{\sigma_z}{\sigma_\theta} = 8$ (c) $\frac{\sigma_z}{\sigma_\theta} = 4$, (d) $\frac{\sigma_z}{\sigma_\theta} = 2$
 (e) $\frac{\sigma_z}{\sigma_\theta} = 1$ (f) $\frac{\sigma_z}{\sigma_\theta} = \frac{1}{2}$.

The ratio of vacuum penetration time through the liner to Alfvén transit time, S , is 100.

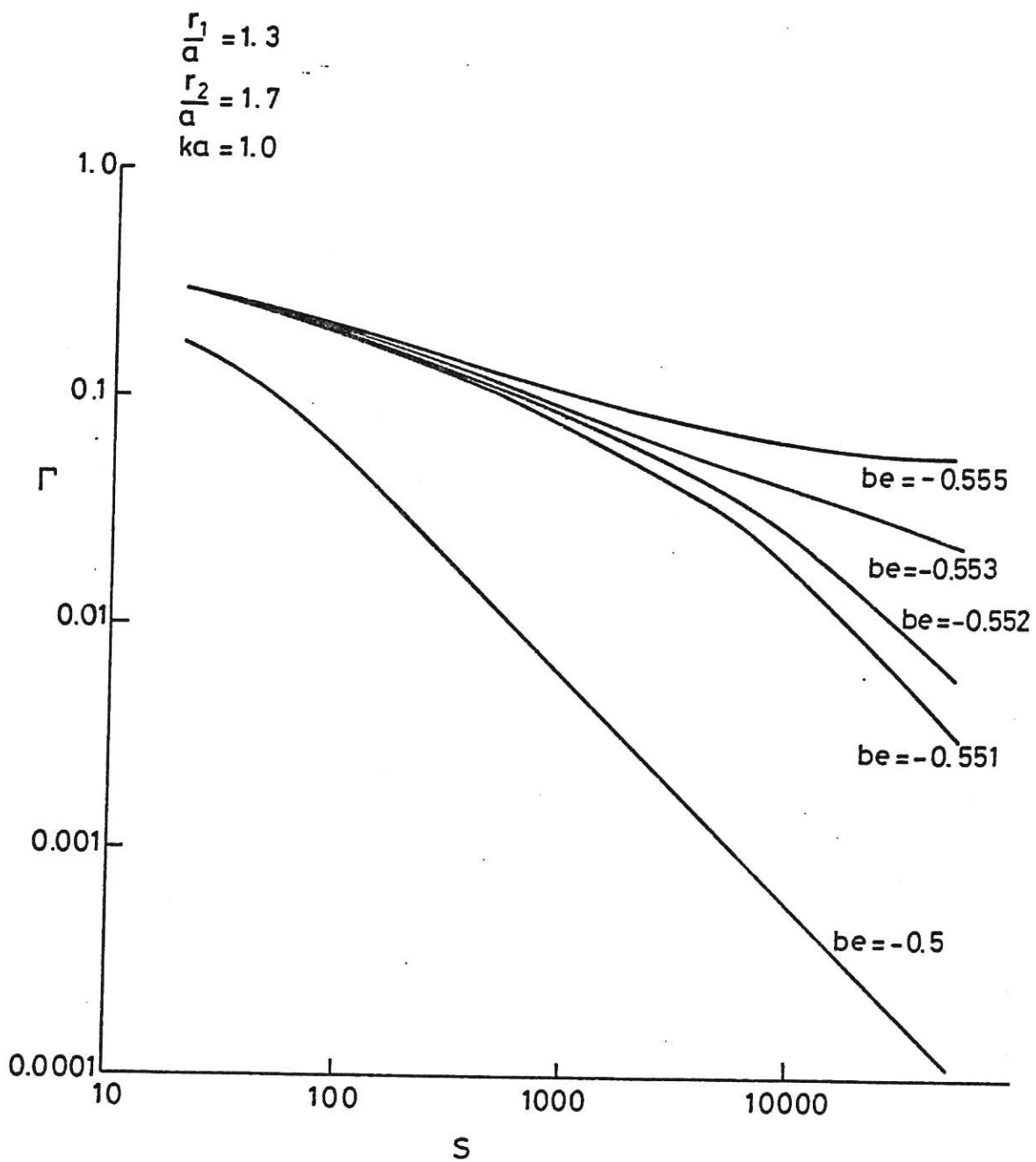


Fig. 3 Normalised growth rate as a function of S as the external longitudinal field, b_e , approaches the marginal stability condition at the liner, $b_e = - .553$.



HER MAJESTY'S STATIONERY OFFICE

Government Bookshops

49 High Holborn, London WC1V 6HB
13a Castle Street, Edinburgh EH2 3AR
41 The Hayes, Cardiff CF1 1JW
Brazennose Street, Manchester M60 8AS
Wine Street, Bristol BS1 2BQ
258 Broad Street, Birmingham B1 2HE
80 Chichester Street, Belfast BT1 4JY

*Government publications are also available
through booksellers*